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Report No. 1

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J. Allen Hynek March 15, 1952

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RESEARCH FOUNDAMES!

Report No....1

REPORT

By

THE OHIO STATE UNIVERSITY RESEARCH FOUNDATION

COLUMBUS 10, OHIO

Cooperator	AP-CAMERIDGE-NESSANCE-LABORATORIES
L40077077 +452740077	230 Albemy Street, Contridge 39, Ness. Contract AF 19(60k)-Al
Investigation of	FLUCTUATIONS OF STARLIGHT AND SKYLLIGHT
Subject of Report	Progress for the period October 1, 1951
4494	to December 31, 1951
Submitted by	J. Allen Hynek
Date Namels 15. 3	052

^{&#}x27;Bubmitted to the Geophysics Research Division, Air Force Cambridge Research Center, Cambridge, Mass. The work reported barein is of a preliminary nature and the results are not necessarily in final form."

ABSTRACT

During this initial period two principal ventures were undertaken:

- (1) the design and construction of a photoelectric photometer employing a 1P21 photomultiplier tube, for use with the 12.5-inch refracting telescope of the McMillin Cheervatory. The photometer features interchangeable filters and disphragms and a solid CO2-cooled tube chamber. The smallest disphragms will admit only one square second of arc of sky, a unique feature of this photometer.
- (2) an extensive survey of the literature, the objective of which is a critical technical report of essentially all known previous work on the problem of astronomical seeing. This report should serve as a critical guide not only for our staff of workers, but for other workers in this field.

The design and construction of the photometer was under the direction of Dr. Geoffrey Escler and Dr. Charles Shaw, with Mr. Protheres in charge of the electronic recording components. The critical survey of the literature is under the direction of Dr. J. A. Hynek, With Mr. Hoger Hosfeld as the chief library research worker.

INVESTIGATIONS IN PROGRESS

(1) Choice and Design of Equipment

Obtaining the observational results that are sought depends on the construction of specialized equipment. The equipment measurery on this project is, first, a telescope of sufficient power to produce an approximate and unasserable transmidisk of a star in the daytime, and second, a device to make a parameter record of the behavior of the image.

The telescope to be used in this work is the 12.5-inch refrector of the McMillin Observatory which will be made available for as many hours as are needed. The telescope is in excellent condition and is operating well. No alterations on the telescope are contemplated, with the possible exception of improving the drive mechanism if this should prove necessary.

The electronic means for recording the fluctuations of the star image, however, must be constructed and essential ab initio, owing to the highly specialized character of the proposed investigation.

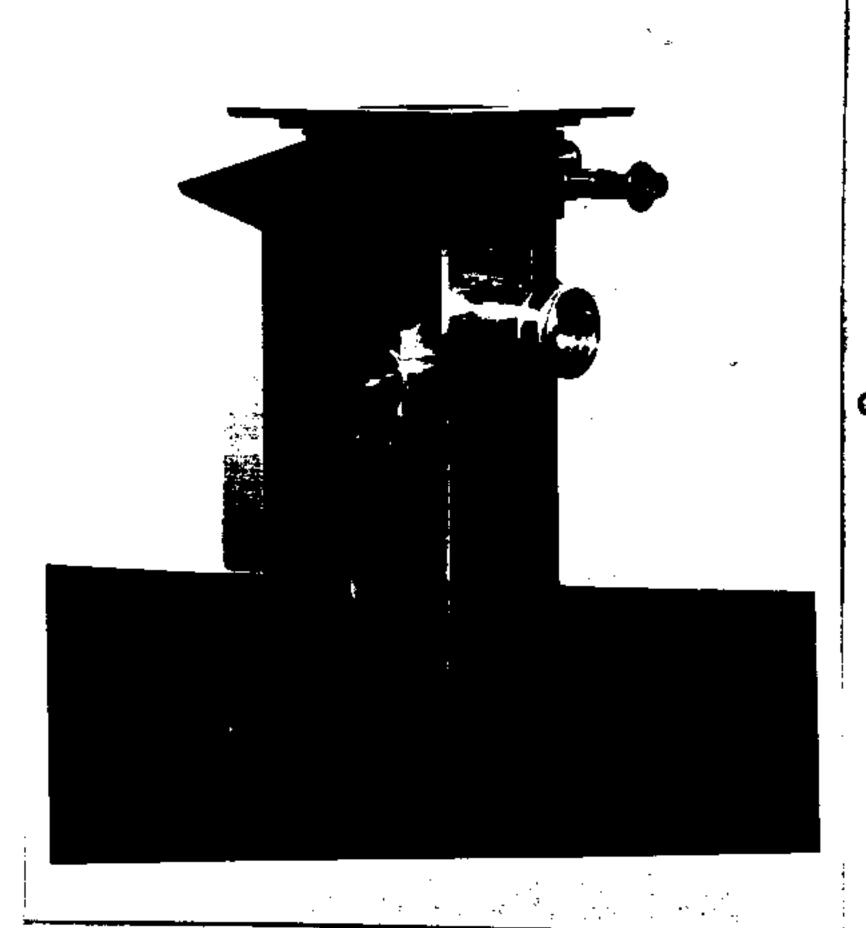
At the initiation of this contract work, therefore, memores concellations were held by Drs. Heeler, Many, Burdle and Hynek, and by Mesers. Protheroe and Hosfeld, among themselves and with other astronomers as to the most appropriate equipment with which to accomplish the project work. It should be pointed out that although Br. Hardle and Mr. Hosfeld were not project staff numbers during this period, they gave freely of their time in the discussions and contributed valuable advice.

Dr. Hynek visited the Mavel Observatory, where similar work is in progress, to consult with Dr. John Hall, director of this work. One of the primary results of this visit, and subsequent correspondence, was a cooperative agreement between the Mavel Observatory and the McMillin Observatory which will eliminate meddens deplication of effort. It was further agreed that there shall be frequent and free interchange of ideas and information. This plan is working sut very well, and Br. Hall and his associates are to be commended for their cooperation and devotion to the principles of the free discussion of scientific results.

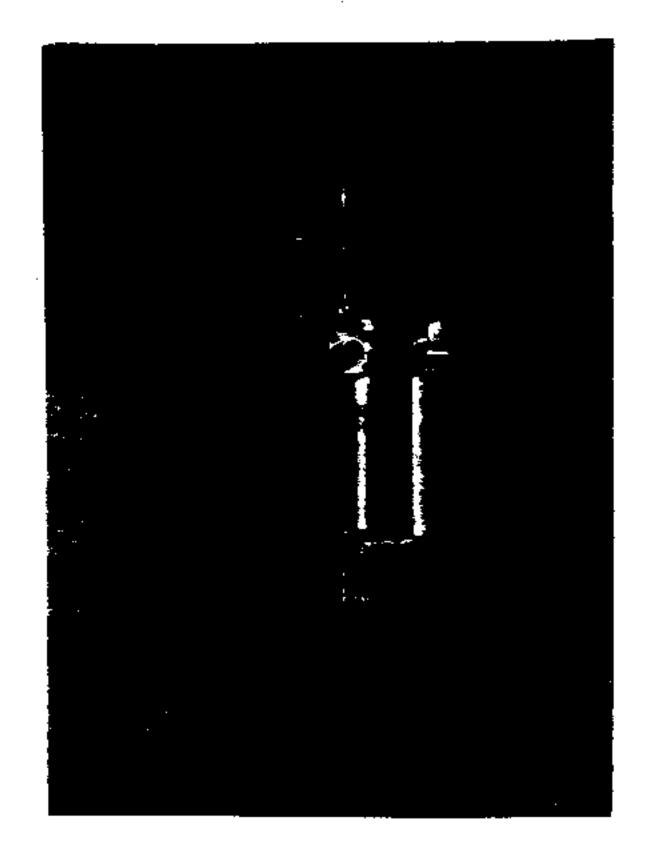
It was agreed that the Mavel Observatory shall continue to specialize in the analysis of the frequency components of stellar scintillation, particularly the very low frequency components, and upon night time astronomical seeing. Our staff, on the other hand, is to concentrate on deptime astronomical seeing, on the quality of the scattered blue daylight sky, and upon the extension of the study of estronomical seeing, both day and night, to the near infrared.

The discussions reported above had a direct bearing on the choice and design of our equipment. It was decided to concentrate on the construction of a highly specialized photoelectric photometer as the primary piece of equipment, to be used directly with the telescope.

The state of construction of the photometer, as of Dec. 31, 1951, is shown in the accompanying photographs. In (a) is shown the optical part of the photometer. The plate shown at the top attaches directly to the double-slide mechanism of the telescope camera housing.



(a) Optical Part of the Photometer



(b) Cooling Chamber for the 1P21 tube Shown immediately below are two eyepieces. The uppermost is a guiding eyepiece. This is moveble over the entire field end is used solely with an auxiliary ster that serves as a "guide star", enabling the observer to insure that the telescope is guiding properly and that, therefore, the ster actually being recorded is playing directly upon the entrance displaying of the optical system.

The wider, aborter eyeploce below the guide eyeploce looks directly into a 45° mirror which can be flipped in and out of place. When the ster to be recorded is on the cross wires of this eyeploce, the observer is assured that when the mirror is flipped ewey (by the hand wheel on the side) the ster is exactly on the entrance disphragm of the spical system.

In view of the extremely small disphragms used (to exclude the maximum amount of daylight sky) it is highly important that the view-ing mirror system be very accurately machined and aligned.

Various disphragms have been mounted on a wheel, controllable from the outside and yet light-tight, so that the variable size of the bramor disk can be properly accompdated.

Immediately behind the entrance disphrage is placed a Fabry less, which images the objective of the telescope (rather than the star) on the cathode of a 1921 photomultiplier.

A wheel, similar to the disphrage wheel but containing a number of Jens and Corning filters, is placed in the optical train before the phototube so that fairly narrow regions of the spectrum (from ultraviolet to infrared) can be edmitted to the cell. (An infrared filter was included even though the 1P21 tube is insensitive to this spectral region, in anticipation of the possible use of a PoS cell.)

Photograph (b) shows the cooling chamber for the 1921 tube. The base of the tube is to be removed, and the loads soldered to brass plates which butt directly against a plug of Tary ice", kept in constant contact with the plates by spring tension. (This part of the design was copied in part from the night time stellar photometer constructed by Dr. Seyfert of the Berneri Chaerystory at Maskville, Tann.)

Standard power supplies food the stages of the IP21 multiplier.

Its output will be recorded on a standard device such as a Brank recorder or a magnetic tape recorder. This part of the electronic train
is essentially standard equipment and need not be elaborated upon here.

(2) Survey of the Literature

At the close of this report period, an intensive survey of the literature had been initiated by Mr. Mosfeld, even though he was not formally on the project until Jan.1, 1952. The Chio State University general library and the astronomical library of the Perkins Sheervatory were used as the principal sources of library material.

Puring this report period the 69-inch reflecting telescope of the Perkins Observatory, the facilities of which are also svailable to our project staff, was used to examine the manner in which starlight illuminates the mirror of this telescope in the daytime. This is done by placing the eye at or near the focus of the telescope and looking out the large mirror. The eye time acts as a Fabry lens and one sees the entire objective illuminated by the starlight.

Long focal lengths are necessary to accomplish the observation of a star-illuminated objective in broad daylight. It can be done with

ease with the 104-foot focal length of the large Perkins reflector, but is virtually impossible with a 15-foot focal length telescope.

We empect to learn a great deal about daytime seeing by a study of the illumination patterns of the 69-inch mirror of the Perkins telescope.

OTHER SCIENTIFIC ACTIVITIES

Conferences

As was noted above, conferences were bold with the Eswal Coservatory staff concerned with similar work.

Correspondence

Copies of two communications from Maval Observatory staff and administrative personnel relative to the agreements entered into with them by Dr. Bynek in the interests of mutual benefit and the elimination of duplication are reproduced herewith (pp. 8 and 9).

It will be noted that overlap is not excluded in these agreements, when it appears profitable to either party. It is the unanimous opinion of our staff that the cooperative arrangement agreed upon is mutually advantageous and that it will enhance without in any way impairing the contract work we are emberked upon.

EXCERPT PROM MEMORAREMENT OF 21 HOYEMBER 1951

09-26 119 Ser:1517726

From: Superintendent, V. S. Bavel Cheervetory

To: Bureau of Ordnessee (Code 329 - Attention Er. H.J. West)

Subj: Butted Project No. 16733

- 1. Recently it came to the attention of the Augerintendent that the Air Force had let a contract to this State University to make investigations under the direction of Dr. Bynek similar to Dated Project No. 16733. The obvious thought occurred that considerable advantage might be gained if these projects were coordinated on a mutually agreeable basis. Discussions of this were held on 16 November 1951 with Dr. Bynek incident to an unsuperted visit of the Boster to the Sheesvatery. Dr. Hall, Mr. Bong and Mr. Milesell attended those discussions for the Chapturkery.
- R. Br. Hyrek is nost desirous of effecting free exchange of information and ideas to svoid meless deglication. He is also agreeable to the establishment of fields of primary effort to be pursued by each of the Charvetories. If the Buyers of Ordanace has no objection, it is the intention of the Superintendent to authorize on a trial basis this conjugative exchange of information and the establishment of fields of primary effort as set forth below:
 - a. Throughout the course of those investigations free exchange of ideas and results in to be maintained.
 - b. The U. S. Barel Chaerrabary will primarily concern itself with (1) the extension of its scintillation measurements to the submulic range, and (2) efforts to determine at what atmospheric levels scintillation occurs.
 - c. The Ohio State Charrestory will be primarily concerned with (1) scintillation measurements in the daytime sky with particular reference to the changes in scintillation with angular distance from the sun, and (2) scintillation measurements in different regions of the spectrum particularly in the infrared.
 - d. No effort will be made to eliminate the everlap expected in the two programs, as an emmination such appears highly desirable.

EXCERPT FROM MEMORANDEM OF 2 DECEMBER 1951

Rega-MJN:gs A1-1

From: Chief, Bureau of Ordunuce

To :: Superintendent U. S. Maval Observatory Washington 25, D. C.

Subj: Bureau of Ordnesce Project No. 16733 and Air Porce contract with Chio State University; coordination of

Ref: (a) USEO ltr Op-26 Al9 Ser 1517P26 of 21 How to BUORD

- 1. The coordination of the work being done by the Eaval Cheervatory on Bureau of Ordnance Project No. 16733 and Ohio State University, under Air Force contract, is satisfactory as outlined in reference (a). It is believed that mutual benefit may result by direct cooperation.
- 2. The Maral Observatory is hereby enthorized to exchange reports with the Ohio State University and provide such assistance as may be practicable and mutually agreeable. It is requested that the Bureau of Ordnance be kept informed of results of this mutual effort by copies of correspondence or such informal reports as may be of interest.
- 3. Br. J. A. Hymak, director of the Air Force project at Chic State University, is informed of this action by copy of this letter and requested to furnish the Bureau of Ordnance, Department of the Mavy, Washington 25, D. C., Attention Re9a, with a copy of reports or results of this study as they become available, in addition to those furnished the U. S. Haval Observatory.

M. F. SCHORYFEL

Copy to: OHIO STATE UNIV Columbus (Dr. J.A. Hynek) E. C. SLEDGE By direction

MARS FOR THE PURCHE

The construction of the photometer should be completed during the next report period, and its testing in the laboratory and on the telescope should be well advanced.

Preliminary observations of very small areas of him sky are also expected to have been made by the close of the next report period (Norch 31, 1952).

The literature survey is expected to be completed and the critical report based on this to be well advanced if not completed.

PERSONNEL AND AMERICAN THE MATTERS

Personnel

The design and construction of the photometer was under the direction of Br. Geoffrey Escler and Br. Charles Shaw, with Mr. Protheros in charge of the electronic recording components. The critical survey of the literature is under the direction of Br. J. A. Hynek, with Mr. Boger Hosfeld as the chief library research worker.

Piscal Information

- (1) The fiscal balance as of Sec. 31, 1951 was \$23,011.76.
 The expenditures during this initial pariod wave showmally small because the project was in the planning and design stage.
- (2) We expect to purchase a brush two-year recorder and double amplifier, a wave emplyment, a suitable tape recorder and purhaps a scangraph. Procurement of the latter item is still highly tempative.

(3) Unassembled electronic and photometer parts are to be procured.

Investigator	
Supervisor J. Allen	Janel Date Mar. 22, 1952
For The Ohio St	ente University Research Foundation Date 24 March 1952